

1 A Simplified Analytical Model for Particle-Geogrid Aperture Interaction

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### Abstract

In this paper, the role of aperture size on the particle-geogrid interaction is analysed based on probabilistic mechanics. Property of ballast size distribution is considered using a Weibull distribution. Particle-geogrid aperture interaction is classified into three mechanisms: interlock interaction, no interaction and sliding interaction, based on the relative particle size that is defined as the ratio of the particle size to the aperture size. Model predictions are in good agreement with a previously proposed correlation between the aperture size and the particle-geogrid interface strength from literature.

### Keywords

geogrids; geosynthetics; theoretical analysis

### List of notations

$\alpha$  interface efficiency factor  
 $\varphi$  friction angle of unreinforced ballast  
 $\delta$  apparent friction angle of geogrid-ballast interface  
 $D$  particle size  
 $A$  aperture size of geogrid  
 $D/A$  relative size  
 $D_{50}$  average particle size  
 $D_{10}$  the particle size that 10% ballast mass is smaller than it  
 $D_{60}$  the particle size that 60% ballast mass is smaller than it  
 $C_u$  the coefficient of uniformity  
 $D_{\min}$  minimal particle size  
 $D_{\max}$  maximum particle size  
 $\alpha(D/A)$  function of  $\alpha$  decided by  $D/A$   
 $f(D)$  probability density for a particle of size  $D$   
 $F(D)$  cumulative probability function of Weibull distribution  
 $P(D)$  probability density function of Weibull distribution  
 $\lambda$  scale parameter of Weibull distribution  
 $k$  shape parameter of Weibull distribution  
 $\xi$   $D/A$  ratio, relative size  
 $P_i$  probability of grain interlock  
 $\alpha_i$  interface efficiency factor of interlock interaction  
 $\alpha_s$  interface efficiency factor of sliding interaction  
 $\alpha_n$  interface efficiency factor when particles do not interlock with a geogrid aperture  
 $d$  standard deviation of normal distribution  
 $h$  a tiny value to integrate probability density

68 **Introduction**

69 Geogrids have been employed to increase the stability of ballasted track in the past few  
70 decades. Several studies have investigated the interaction of the aperture size ( $A$ ) of geogrids  
71 and ballast (McDowell et al., 2006; Brown et al., 2007; Indraratna et al., 2011; Sweta and  
72 Hussaini, 2018). McDowell et al. (2006) reported the optimum ratio between aperture size and  
73 nominal particle size to be in the proximity of 1.4 in relation to the peak resistance in a discrete  
74 element simulation of pull-out testing. Brown et al. (2007) concluded that the optimum aperture  
75 size for ballast with an average particle size of 50mm should be 60-80mm to achieve the lowest  
76 settlement after 3000 cycles of loading. Hussaini (2012) identified the role of aperture size on  
77 the reduction of lateral displacement of geogrid-reinforced ballast interface under cyclic loading  
78 and reported the optimum  $A/D_{50}$  ratio to be 1.21, where  $D_{50}$  is the average particle size.  
79 Indraratna et al. (2014) reported the optimum  $A/D_{50}$  ratio in the range of 1.1 for the least  
80 geogrid-reinforced ballast deformation measured in the field. Palmeira and G3ngora (2016)  
81 conducted several large scale cyclic loading tests for geogrid-reinforced gravel fill and reported  
82 the effect of aperture size on the traffic benefit ratio (TBR), which is defined as the number of  
83 loading cycles of reinforced material to that of unreinforced material for the same rut depth at  
84 surface. They concluded the optimum  $A/D_{50}$  ratio to be approximately equal to 2. Indraratna et  
85 al. (2011) reported a correlation between aperture size ( $A$ ) and interface efficiency factor ( $\alpha$ )  
86 (Koerner 1998), defined as:

87

88 
$$\alpha = \tan \delta / \tan \varphi \tag{1}$$

89

90 where  $\varphi$  is the friction angle of unreinforced ballast and  $\delta$  is the apparent friction angle at the  
91 interface of geogrid and ballast. And the optimum aperture size was reported as  $1.2D_{50}$ . Sweta  
92 and Hussaini (2018) identified the similar correlations between  $\alpha$  and  $A/D_{50}$  under different  
93 shearing rates.

94

95 In terms of theoretical modelling, Jewell et al. (1984) derived a micromechanical framework for  
96 the geogrid-reinforced soil. Greenwood and Williamson (1966) developed a model for the  
97 interaction between two rough surfaces. However, the interaction between geogrid and ballast is

98 different with above interactions, as ballast particles have various sizes resulting in different  
99 interaction mechanisms with geogrid. Therefore, this paper aims to develop an analysis to  
100 capture and quantify the relationship between the aperture size ( $A$ ) and interface efficiency  
101 factor ( $\alpha$ ) using the particle size distribution.

102

## 103 **2. Theoretical Considerations**

### 104 **2.1 Interface efficiency factor**

105 For simplicity, this study assumes that particle and aperture size has major influence on the  
106 interface efficiency factor, compared to other properties of geogrids such as its thickness,  
107 tensile strength, bending stiffness of transverse members, junction strength etc. The particle-  
108 aperture interaction effect is assumed to be primarily dependent on the relative particle size,  
109 defined as the ratio of the particle size ( $D$ ) to the aperture size of geogrid ( $A$ ), denoted as  $D/A$ .  
110 Thus, the interface efficiency factor ( $\alpha$ ) can be taken as the average of the interface efficiency  
111 factors of all particles, hence:

112

$$113 \quad \alpha = \int_{D_{min}}^{D_{max}} \alpha(D/A) f(D) dD \quad (2)$$

114

115 where  $D_{min}$  and  $D_{max}$  are the minimum and maximum particle size of ballast, respectively,  
116  $\alpha(D/A)$  is the function of  $\alpha$  controlled by  $D/A$ ,  $f(D)$  is the probability density for a particle of size  
117  $D$  at the interface. For a given  $f(D)$ ,  $\alpha$  is the function of  $A$ . In contrast to the empirical relationship  
118 where the interface efficiency factor is characterized using mean ballast size ( $D_{50}$ ), Equation (2)  
119 considers the entire particle size distribution.

120

### 121 **2.2 Evaluation of $f(D)$**

122  $f(D)$  can be derived from particle size distribution (PSD) curve. In this Note, Weibull distribution  
123 (Weibull, 1951) is adopted to describe ballast PSD (Fang et al, 1993). The cumulative  
124 distribution function (CDF) of Weibull distribution is given by:

125

$$126 \quad F(D) = 1 - e^{-\left(\frac{D}{\lambda}\right)^k} \quad (3)$$

127

128 The probability density function (PDF) is presented by:

129

$$130 \quad P(D) = \frac{k}{\lambda} \left(\frac{D}{\lambda}\right)^{k-1} e^{-(D/\lambda)^k} \quad (4)$$

131

132 where  $\lambda$  is the scale parameter and  $k$  is the shape parameter. The influences of  $\lambda$  and  $k$  on CDF

133 and PDF are shown in Figure 1. The PDF of Weibull Distribution can be taken as  $f(D)$ . The

134 correlation between  $\lambda$ ,  $k$  and PSD parameters can be derived as follow. Taken  $D$  as  $\lambda$ , Equation

135 (3) can be derived into:

136

$$137 \quad F(\lambda) = 1 - e^{-\left(\frac{\lambda}{\lambda}\right)^k} = 1 - e^{-1} \approx 0.6 \quad (5)$$

138

139 Therefore, the probability of  $D \leq \lambda$  is approximately 0.6, and the physical meaning of  $\lambda$  is the

140 particle size at which 60% ballast mass is smaller than, namely  $D_{60}$ . Taken  $D$  as  $D_{10}$ , where  $D_{10}$

141 is the particle size at which 10% ballast mass is smaller than, then Equation (3) can be re-

142 written as:

143

$$144 \quad F(D_{10}) = 1 - e^{-\left(\frac{D_{10}}{\lambda}\right)^k} \quad (6)$$

145

146 Recalling the physical meaning of  $\lambda$  and  $D_{10}$ , Equation (6) becomes:

147

$$148 \quad 0.1 = 1 - e^{-\left(\frac{D_{10}}{D_{60}}\right)^k} = 1 - e^{-\left(\frac{1}{C_u}\right)^k} \quad (7)$$

149

150 where,  $C_u$  is the coefficient of uniformity. Therefore, the correlation between  $k$  and  $C_u$  can be

151 written as:

152

$$153 \quad k = \frac{1}{\log(C_u)} \quad (8)$$

154

155 **2.3 Evaluation of  $\alpha(D/A)$**

156 Following several past studies (Indraratna et al., 2011; Sweta and Hussaini, 2018 and Han et  
157 al., 2018), the role of aperture size on the shear strength of particle-geogrid interface can be  
158 commonly attributed to three distinct interaction mechanisms:

159

160 (1) Interlock Interaction: Once the interlock between geogrids and particles is formed,  $\alpha$  is  
161 assumed to have a constant value,  $\alpha_i$ .

162

163 (2) No Interaction: When  $D < A$ , unless a cluster of particles is considered to be interlocking  
164 with a given geogrid aperture, single particles are too small to interact with the geogrid. Thus,  $\alpha$   
165 = 1 when  $D < A$ .

166

167 (3) Sliding Interaction: When  $D \geq A$ , if particles do not interlock geogrids with their smaller  
168 dimensions, particles are considered to be too large to form a proper interlock, thus the particle  
169 tends to slide along the grid interface. In this case,  $\alpha$  is assumed to have the constant value,  $\alpha_s$ .  
170 According to lab observations of Indraratna et al. (2011), Sweta and Hussaini, (2018) and Han  
171 et al. (2018), interface efficiency factor is smaller than unity when aperture size is relatively  
172 smaller than particle size. If aperture size is extremely small, particle-geogrid interaction would  
173 work like particle-geotextile interaction. Therefore,  $\alpha_s$  should be smaller than 1 but larger than  
174 the interface efficiency factor of ballast-geotextile interaction, which is 0.8 based on the  
175 laboratory measurement for ballast-geotextile interaction as reported in an earlier study by  
176 Indraratna et al. (2011). To clarify further the evaluation of  $\alpha(D/A)$ , a dimensionless variable  $\xi$  is  
177 adopted to replace  $D/A$ . Therefore,  $\alpha(D/A)$  can be expressed as:

178

179 
$$\alpha(D/A) = \alpha(\xi) = P_i(\xi) \times \alpha_i + (1 - P_i(\xi)) \times \alpha_n \quad (9)$$

180

181 where  $P_i(\xi)$  is the probability of grain interlock for a specific  $D/A$  ratio ( $\xi$ ),  $\alpha_n$  is the interface  
182 efficiency factor when particles do not interlock with a geogrid aperture, where  $\alpha_n = 1$  when  $\xi <$   
183  $1$ , and  $\alpha_n = \alpha_s$  when  $\xi \geq 1$ . As the discontinuity of  $\alpha_n$  at  $\xi = 1$  seldom occurs in practice, the  
184 following equation is adopted to describe  $\alpha_n$  as also plotted in Figure 2.

185

$$186 \quad \alpha_n(\xi) = \alpha_s + (1 - \alpha_s)/(1 + e^{20 \times (\xi - 1)}) \quad (10)$$

187

188 The probability density of grain interlock is assumed to be a normal distribution whose mean  
189 equals to 1, and when  $\xi = 1$ , the highest probability of sustaining a particle-geogrid interlock is  
190 achieved. Therefore,  $P_i(\xi)$  can be expressed as follows;

191

$$192 \quad P_i(\xi) = \int_{\xi-h}^{\xi+h} \frac{1}{d\sqrt{2\pi}} e^{-\frac{1}{2}\left(\frac{\xi-1}{d}\right)^2} \quad (11)$$

193

194 where  $d$  is the standard deviation of normal distribution,  $h$  is a tiny value to integrate probability  
195 density.

196

197 To determine values of  $\alpha_i$ ,  $\alpha_s$ ,  $d$  and  $h$  discussed above, a series of direct shear tests was  
198 conducted using ballast with uniform particle size and geogrids under the same procedures  
199 suggested by Indraratna et al. (2011). Single-size particles were sieved and divided into four  
200 size intervals: 13.2-19mm, 19-26.5mm, 26.5-37.5mm, 37.5-53mm. Uniform sizes of ballast were  
201 established by averaging two sizes of intermittent sieves as 16mm, 23mm, 32mm and 45mm.  
202 Physical characteristics of geogrids used in testing are shown in Table 1. The equivalent  
203 aperture sizes (21-64mm) are derived according to the approach explained by Indraratna et al.  
204 (2011), which is the square root of the opening area that is taken as the equivalent aperture size  
205 of rectangular apertures. The diameter of the largest inscribed circle is taken as the equivalent  
206 aperture size of triangular apertures. The interface efficiency factors ( $\alpha$ ) for various A/D ratios ( $\xi$ )  
207 are shown in Table 2. Experimental results are plotted in Figure 3.

208

209 When  $\alpha_i = 2.9$ ,  $\alpha_s = 0.83$ ,  $d = 0.17$  and  $h = 0.05$ , Equation (9) has a reasonable fit with  
210 experimental data. With obtained values of  $d$  and  $h$ , Equation (11) is plotted in Figure 4. As  
211 shown in Figure 4, the probability of grain interlock reaches the maximum of 23.1% when  $\xi = 1$ ,  
212 and it decreases to zero when  $\xi = 0.5$  and 1.5. This means that small particles whose size is

213 smaller than 0.5A and large particles whose size is larger than 1.5A have extremely low  
214 probability of interlocking with the geogrid apertures.

215

### 216 **3. Model Validation**

217 To capture the effect of entire PSD, a set of large-scale direct shear test data from Indraratna et  
218 al. (2011) and from laboratory investigations using similar procedures by the Authors are used  
219 to validate the proposed model. The fresh latite ballast PSDs used in the current study and by  
220 Indraratna et al. (2011) are shown in Figure 5. Additional direct shear test on well-graded ballast  
221 with the coefficient of uniformity of 2.3 was conducted in this study. It is noted that Indraratna et  
222 al. (2011) adopted a more uniform ballast gradation with a coefficient of uniformity of 1.87.  
223 Physical characteristics of geogrids used in testing are shown in Table 1. Corresponding  
224 interface efficiency factors are shown in Table 3. The parameters used in model validation are  
225 listed in Table 4.

226

227 The comparison between model predictions and the experimental data from Indraratna et al.  
228 (2011) and current study is shown in Figure 6. It shows a good agreement between the  
229 theoretical predictions and the experimental data. It can be found that the variation of the  
230 interface friction of more uniform ballast conducted by Indraratna et al. (2011) has a slightly  
231 higher peak of  $\alpha$  with a narrower range of aperture sizes having  $\alpha > 1$  when compared to that of  
232 less uniform ballast used in the current study. The reinforcing effect is provided purely by the  
233 interlock between particle and geogrids. The probability of grain interlock for specific PSD can  
234 be evaluated as follow

235

$$236 \quad P_i = \int_{D_{min}}^{D_{max}} P_i(\xi) f(D) dD \quad (12)$$

237

238 When  $A/D_{50} = 1.35$ , the probabilities of grain interlock are 11.2% for more uniform PSD, and  
239 9.3% for more well-graded PSD. The higher probability of grain interlock for more uniform PSD  
240 makes  $\alpha$  of Indraratna et al. (2011) higher than that of the current study. However, when  $A/D_{50}$   
241 = 2.2, the probabilities of grain interlock are 1.6% and 1.7% for more uniform and more well-  
242 graded PSD, respectively, resulting in a lower value of  $\alpha$  for more uniform ballast. This model



243 shows that different PSD gradations have different probabilities of grain interlock, leading to  
244 different interface behaviours. It can be seen that the model can reasonably capture the  
245 aperture size effect on different ballast gradations.

246

#### 247 **4. Model Limitations**

248 This model considers the interaction between particles and apertures of a geogrid, hence it has  
249 limitations as explained below. Apart from the relative size of particle in relation to grid aperture  
250 size, other properties of geogrids such as its thickness, tensile strength, bending stiffness of  
251 transverse members, junction strength etc., can influence the grid-ballast interaction (Brown et  
252 al., 2007; Cuelho et al., 2014; Palmeira and Góngora, 2016). However, these parameters of  
253 geogrids have been excluded from the current mathematical model for simplicity.

254

#### 255 **5. Conclusions**

256 In this study, the use of probabilistic analysis in predicting the correlation between aperture size  
257 ( $A$ ) and the interface efficiency factor ( $\alpha$ ) by exploiting three distinct interaction mechanisms is  
258 demonstrated. The proposed three particle-geogrid interaction mechanisms, namely: interlock  
259 interaction, no interaction and sliding interaction, are consistent with the findings of laboratory  
260 investigations from other researchers. The developed analytical model considers the ballast size  
261 gradation using Weibull distribution.

262

263 The developed theoretical model is particularly advantageous when ballast size gradations have  
264 different coefficients of uniformity ( $C_u$ ), which are not reflected by the  $D_{50}$ . The resulting model  
265 predictions are in good agreement with the interface efficiency factors ( $\alpha$ ) based on two sets of  
266 laboratory studies conducted by the authors and through past literature. When the particle size  
267 distribution becomes more uniform, the interface efficiency factor increases at the optimum  
268 aperture size but decreases at non-optimum aperture sizes. The model can represent well the  
269 variation of the interface behaviour between ballast and geogrids considering the geogrid's  
270 aperture size and particle size distribution, which previous models failed to capture.

271

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**Table 1 Physical characteristics of the geogrids used in the current study**

Geogrids	Type	Aperture Shape	Aperture Size, A (mm)	Rib Width (mm)	Ultimate Tensile
					Strength (kN/m)
G1	Extruded	Triangle	21*	1	19
G2	Woven	Square	25	6	43
G3	Welded	Square	32	7	30
G4	Woven	Rectangle	43*	8	43
G5	Extruded	Rectangle	64*	5	30

334 Note: \* Equivalent aperture size; All geogrids are biaxial products

**Table 2 Interface efficiency factors for various D/A ratios ( $\xi$ )**

Geogrids Type	Uniform Ballast Size, D (mm)	D/A ratios, $\xi$	Interface Efficiency Factor, $\alpha$
G1	23	1.09	1.15
	32	1.52	0.88
G2	23	0.92	1.26
	32	1.28	1.05
G3	16	0.5	1.00
	23	0.72	1.07
	32	1	1.33
	45	1.41	0.89
G4	23	0.53	1.01
	32	0.74	1.04
G5	23	0.36	1.00
	32	0.5	1.00

336 **Table 3 Interface efficiency factors for well-graded PSD ( $C_u = 2.3$ ) measured in current**  
 337 **study**

Geogrids	Interface Efficiency Factor, $\alpha$
G1	0.93
G2	0.94
G3	1.11
G4	1.16
G5	1.05

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339

340

**Table 4 Model Parameters**

Data Source	$\alpha_i$	$\alpha_s$	d	h	$D_{min}$	$D_{max}$	$\lambda$	k
Current Study	2.9	0.83	0.17	0.05	5	60	38	2.75
Indraratna et al. (2011)					10	60	38	3.68

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356 **Figure captions**

357 **Figure 1.** (a) Weibull cumulative distribution function, Equation (3); (b) Weibull probability  
358 density function, Equation (4)

359 **Figure 2.** Correlation between  $\alpha_n$  and  $\xi$  according to Equation (10)

360 **Figure 3.** Comparisons between Equation (9) and lab data when  $\alpha_i = 2.9$ ,  $\alpha_s = 0.83$ ,  $d = 0.17$   
361 and  $h = 0.05$

362 **Figure 4.** Plot of Equation (11) when  $d = 0.17$  and  $h = 0.05$

363 **Figure 5.** Particle size distributions used in current study and Indraratna et al. (2011)

364 **Figure 6.** Comparisons between model predictions and experimental results

365